# Repulsive gravity model for dark energy

#### Manuel Hohmann

II. Institut für theoretische Physik



DPG-Tagung Karlsruhe 2011

#### • ACDM model: only 5% visible matter.

- Dark matter explains "missing mass" in galaxies.
- Dark energy explains accelerating expansion.
- Constituents of dark universe are unknown!

- ACDM model: only 5% visible matter.
  - Dark matter explains "missing mass" in galaxies.
  - Dark energy explains accelerating expansion.
  - Constituents of dark universe are unknown!
- Idea here: Additional "dark, negative mass" standard model copy.
- Only interaction between both copies: repulsive gravity.
- Universe contains equal amounts of both types of mass.

- ACDM model: only 5% visible matter.
  - Dark matter explains "missing mass" in galaxies.
  - Dark energy explains accelerating expansion.
  - Constituents of dark universe are unknown!
- Idea here: Additional "dark, negative mass" standard model copy.
- Only interaction between both copies: repulsive gravity.
- Universe contains equal amounts of both types of mass.
- Dark galaxies "push" visible matter & light towards visible galaxies.
   ⇒ Explanation of dark matter!
- Mutual repulsion between galaxies drives accelerating expansion.
   ⇒ Explanation of dark energy!

- Positive and negative test masses follow different trajectories.
- Two types of test mass trajectories  $\Rightarrow$  two types of observers.
- Observer trajectories are autoparallels of two connections ∇<sup>±</sup>.
- Observers attach parallely transported frames to their curves.
- Frames are orthonormalized using two metric tensors g<sup>±</sup><sub>ab</sub>.

- Positive and negative test masses follow different trajectories.
- Two types of test mass trajectories  $\Rightarrow$  two types of observers.
- Observer trajectories are autoparallels of two connections ∇<sup>±</sup>.
- Observers attach parallely transported frames to their curves.
- Frames are orthonormalized using two metric tensors  $g_{ab}^{\pm}$ .
- No-go theorem forbids bimetric repulsive gravity. [MH, M. Wohlfarth '09]
- Solution:  $N \ge 3$  metrics  $g_{ab}^{l}$  and standard model copies  $\Psi^{l}$ .

# Action and equations of motion

- Matter action: sum of standard model actions.
- Gravitational action:

$$S_G[g^1, ..., g^N] = rac{1}{2} \int d^4 x \sqrt{g_0} \left[ \sum_{l,J=1}^N (x + y \delta^{lJ}) g^{lij} R^J_{ij} + F(S^{lJ}) 
ight]$$

- Symmetric volume form  $g_0 = (g^1 g^2 \dots g^N)^{1/N}$ .
- $F(S^{IJ})$  quadratic in connection difference tensors  $S^{IJ} = \Gamma^{I} \Gamma^{J}$ .

# Action and equations of motion

- Matter action: sum of standard model actions.
- Gravitational action:

$$S_G[g^1, \dots, g^N] = rac{1}{2} \int d^4 x \sqrt{g_0} \left[ \sum_{I,J=1}^N (x + y \delta^{IJ}) g^{Iij} R^J_{ij} + F(S^{IJ}) 
ight]$$

• Symmetric volume form  $g_0 = (g^1 g^2 \dots g^N)^{1/N}$ .

•  $F(S^{IJ})$  quadratic in connection difference tensors  $S^{IJ} = \Gamma^{I} - \Gamma^{J}$ .  $\Rightarrow$  Equations of motion:

$$T_{ab}^{I} = \sqrt{\frac{g_0}{g^{I}}} \left[ -\frac{1}{2N} g_{ab}^{I} \sum_{J,K=1}^{N} (x + y\delta^{JK}) g^{Jij} R_{ij}^{K} + \sum_{J=1}^{N} (x + y\delta^{JJ}) R_{ab}^{J} \right]$$

+ terms linear in  $\nabla^{I} S^{JK}$ 

+ terms quadratic in  $S^{IJ}$ .

#### ⇒ Repulsive Newtonian limit for $N \ge 3$ .

# Simple cosmological model

- Homogeneous and isotropic universe (FLRW metric).
- Matter content: perfect fluid.
- Early universe: radiation; late universe: dust.
- Copernican principle: common evolution for all matter sectors.

# Simple cosmological model

- Homogeneous and isotropic universe (FLRW metric).
- Matter content: perfect fluid.
- Early universe: radiation; late universe: dust.
- Copernican principle: common evolution for all matter sectors.
- $\Rightarrow$  Single effective energy-momentum tensor  $T'_{ab} = T_{ab}$ .
- $\Rightarrow$  Single effective metric  $g_{ab}^{\prime} = g_{ab}$ .
- $\Rightarrow$  Ricci tensors  $R_{ab}^{l} = R_{ab}$  become equal.
- $\Rightarrow$  Connection differences  $S^{IJi}_{jk} = 0$  vanish.

# Simple cosmological model

- Homogeneous and isotropic universe (FLRW metric).
- Matter content: perfect fluid.
- Early universe: radiation; late universe: dust.
- Copernican principle: common evolution for all matter sectors.
- $\Rightarrow$  Single effective energy-momentum tensor  $T'_{ab} = T_{ab}$ .
- $\Rightarrow$  Single effective metric  $g_{ab}^{\prime} = g_{ab}$ .
- $\Rightarrow$  Ricci tensors  $R_{ab}^{l} = R_{ab}$  become equal.
- $\Rightarrow$  Connection differences  $S^{IJi}_{jk} = 0$  vanish.
- $\Rightarrow$  Equations of motion simplify for repulsive Newtonian limit:

$$(2-N)T_{ab}=R_{ab}-\frac{1}{2}Rg_{ab}.$$

 $\Rightarrow$  Negative effective gravitational constant for early / late universe.

# Multimetric cosmology

• Derive cosmological equations of motion:

$$\frac{\ddot{a}}{a}=rac{N-2}{6}\left(
ho+3
ho
ight).$$

 $\Rightarrow$  Acceleration must be positive.

# Multimetric cosmology

Derive cosmological equations of motion:

$$\frac{\ddot{a}}{a}=\frac{N-2}{6}\left(\rho+3\rho\right).$$

- $\Rightarrow$  Acceleration must be positive.
  - Explicit solution for radiation (dashed) / dust (solid):



#### ⇒ Big bounce. [MH, M. Wohlfarth '10]

Manuel Hohmann (Uni Hamburg)



• *N* = 4.

 Different matter types separate.

#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)



• *N* = 4.

 Different matter types separate.

#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)



#### • *N* = 4.

 Different matter types separate.

#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)



• *N* = 4.

 Different matter types separate.

#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)



Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)

matter types separate.

• N=4. Different

•



• *N* = 4.

 Different matter types separate.

#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)



• *N* = 4.

 Different matter types separate.

#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)



• *N* = 4.

 Different matter types separate.

#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)



• *N* = 4.

 Different matter types separate.

#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)



• *N* = 4.

 Different matter types separate.

#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)



• *N* = 4.

 Different matter types separate.

#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)



• *N* = 4.

 Different matter types separate.

#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)



• *N* = 4.

 Different matter types separate.

#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)



• *N* = 4.

 Different matter types separate.

#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)



• *N* = 4.

 Different matter types separate.

#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)



• *N* = 4.

 Different matter types separate.

#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)



• *N* = 4.

 Different matter types separate.

#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)



• *N* = 4.

 Different matter types separate.

#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)



• *N* = 4.

 Different matter types separate.

#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)



• *N* = 4.

 Different matter types separate.

#### Video: www.desy.de/ mhohmann



• *N* = 4.

 Different matter types separate.

#### Video: www.desy.de/ mhohmann

- Galactic clusters.
- Empty voids.



#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)

- Galactic clusters.
- Empty voids.



#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)

- Galactic clusters.
- Empty voids.



#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)

- Galactic clusters.
- Empty voids.



#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)

- Galactic clusters.
- Empty voids.



#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)

- Galactic clusters.
- Empty voids.



#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)

- Galactic clusters.
- Empty voids.



#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)

- Galactic clusters.
- Empty voids.



#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)

- Galactic clusters.
- Empty voids.



#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)

- Galactic clusters.
- Empty voids.



#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)

- Galactic clusters.
- Empty voids.



#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)

- Galactic clusters.
- Empty voids.

#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)

#### Galactic clusters.

Empty voids.

#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)

- Galactic clusters.
- Empty voids.



#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)

- Galactic clusters.
- Empty voids.



#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)

- Galactic clusters.
- Empty voids.



#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)

#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)

Galactic

clusters.Empty voids.

- Galactic clusters.
- Empty voids.



#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)

- Galactic clusters.
- Empty voids.



#### Video: www.desy.de/ mhohmann

Manuel Hohmann (Uni Hamburg)

- Galactic clusters.
- Empty voids.

#### Video: www.desy.de/ mhohmann



- Galactic clusters.
- Empty voids.

Video: www.desy.de/ mhohmann

# Parametrized post-Newtonian formalism

- Obtain "fingerprint" of single-metric gravity theories. [Thorne, Will '71; Will '93]
- $\Rightarrow$  10 parameters, constrained by solar system experiments.

# Parametrized post-Newtonian formalism

- Obtain "fingerprint" of single-metric gravity theories. [Thorne, Will '71; Will '93]
- $\Rightarrow$  10 parameters, constrained by solar system experiments.
- Extension to multimetric gravity theories. [MH, M. Wohlfarth '10]
- $\Rightarrow$  Additional 14 unobserved parameters.
- $\Rightarrow$  8 parameters can be obtained from linearized field equations.

# Parametrized post-Newtonian formalism

- Obtain "fingerprint" of single-metric gravity theories. [Thorne, Will '71; Will '93]
- $\Rightarrow$  10 parameters, constrained by solar system experiments.
  - Extension to multimetric gravity theories. [MH, M. Wohlfarth '10]
- $\Rightarrow$  Additional 14 unobserved parameters.
- $\Rightarrow$  8 parameters can be obtained from linearized field equations.
- Example: action can be chosen such that
  - $\alpha^+ = 1$ ,  $\theta^+ = 0$ : standard PPN gauge choice.
  - $\gamma^+ = 1$ ,  $\sigma^+_+ = -2$ : experimental consistency.
  - $\alpha^- = -1$ : repulsive Newtonian limit.
  - $\gamma^{-} = -1$ ,  $\theta^{-} = 0$ ,  $\sigma^{-}_{+} = 2$ : additional "dark" PPN parameters.

 $\Rightarrow$  Consistent with solar system experiments up to linear PPN order.

# Gravitational waves

- Propagation velocity equals speed of light.
- Up to 6 polarizations in general metric theories.
- Theories classified by representations of E(2).
- E(2) class of multimetric gravity depends on 3 parameters:



- PPN consistent theory shown before of class N<sub>2</sub>.
- $\Rightarrow$  Same class as general relativity.

- Idea: Repulsive gravity might explain dark matter & dark energy.
- $\Rightarrow$  General relativity must be extended to allow repulsive gravity.
- $\Rightarrow$  No-go theorem: bimetric repulsive gravity is not possible.
- ⇒ Multimetric repulsive gravity with  $N \ge 3$  by explicit construction.
- $\Rightarrow$  Cosmology features late-time acceleration and big bounce.
- $\Rightarrow$  Structure formation features clusters and voids.
- $\Rightarrow$  Repulsive gravity is consistent with solar system experiments.
- $\Rightarrow$  Gravitational waves are null.
- $\Rightarrow~$  E(2) class can be one of  $\rm N_2,\,N_3,\,III_5,\,II_6.$

# Outlook

- Remaining PPN parameters should be determined from full multimetric PPN formalism.
- Restrict multimetric gravity theories by additional PPN bounds.
- Establish further construction principles, e.g., continuous symmetry between sectors.
- Examine initial-value problem.
- Determine further exact solutions (single point mass...).
- Advanced simulation of structure formation including thermodynamics using GADGET-2 (Millenium Simulation).
- Search for repulsive gravity sources in the galactic voids through gravitational lensing.
- Application to binaries: gravitational radiation should be emitted in all sectors, but only one type is visible.

# Outlook

- Remaining PPN parameters should be determined from full multimetric PPN formalism.
- Restrict multimetric gravity theories by additional PPN bounds.
- Establish further construction principles, e.g., continuous symmetry between sectors.
- Examine initial-value problem.
- Determine further exact solutions (single point mass...).
- Advanced simulation of structure formation including thermodynamics using GADGET-2 (Millenium Simulation).
- Search for repulsive gravity sources in the galactic voids through gravitational lensing. Prediction!
- Application to binaries: gravitational radiation should be emitted in all sectors, but only one type is visible. Prediction!