

The Dawn of Black Hole Astronomy

What can be inferred from the image of M87*? What are the next tasks and targets?

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Center of Excellence "The Dark Side of the Universe"



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How does the size of a black hole depend on its mass?

- Ordinary matter:
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THE M87 BLACK HOLE
AND
OUR SOLAR SYSTEM



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⇒ Support for linear relation.

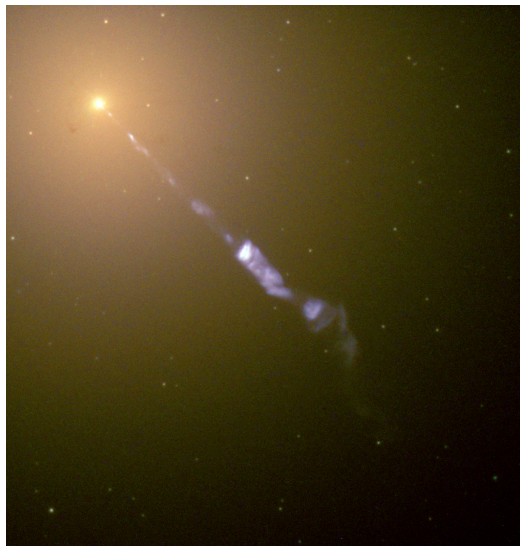
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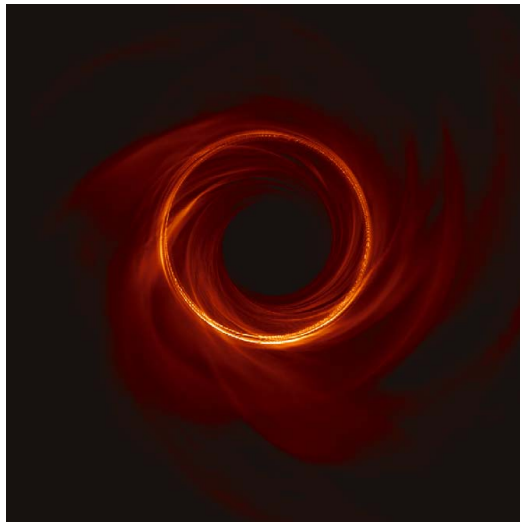
Does M87* spin? A look at its jet and accretion disk

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 - Extends $\approx 1.5\text{kpc}$ from the center.
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 - Collimation hints towards magnetic fields.
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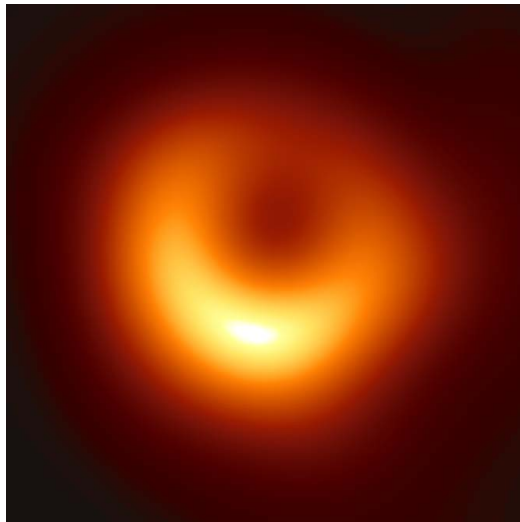
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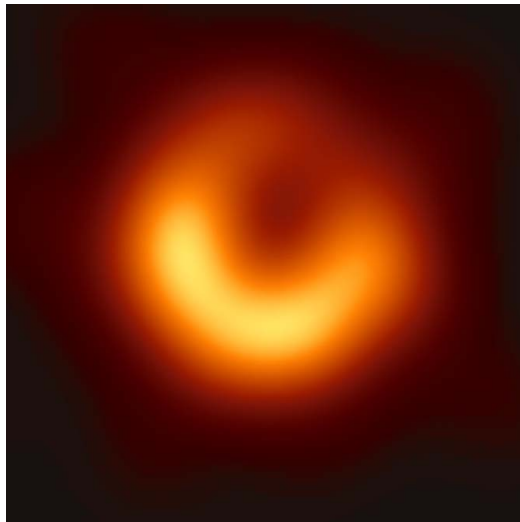
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 - ⇒ **Consistent with spinning central object.**



Is M87* really a black hole, and possess a horizon?

- What the Event Horizon Telescope does *not* observe:
 - The EHT does *not* (directly) observe event horizons of black holes.
 - There is no direct observable linked to the existence of a horizon.
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 - What can we learn about the nature of the central object in M87*?
 - Some types of horizonless objects may allow radiation to...
 - pass through (boson star, wormhole...),
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- ⇒ EHT observation shows dark central area, no signs of traversing or reflected radiation.
- The accretion disk around a black hole (and similar objects) is truncated at the ISCO.
- ⇒ No visible inflow within the central region, consistent with a truncated disk.

Getting more information by including polarizations

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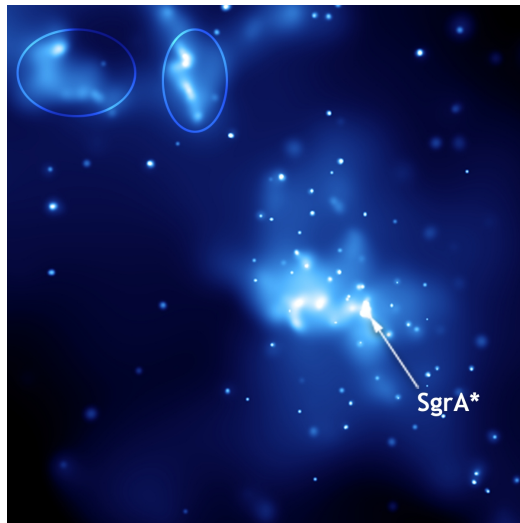
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- Can polarization tell us anything about gravity and the central object?
 - In general relativity, light propagates independent of polarization in vacuum.
 - Alternative theories may include “vacuum birefringence” - speed of light differs.
 - Lensing at a massive object may lead to two polarized images of what is behind.
 - ⇒ Observational signature to seek in future observations.

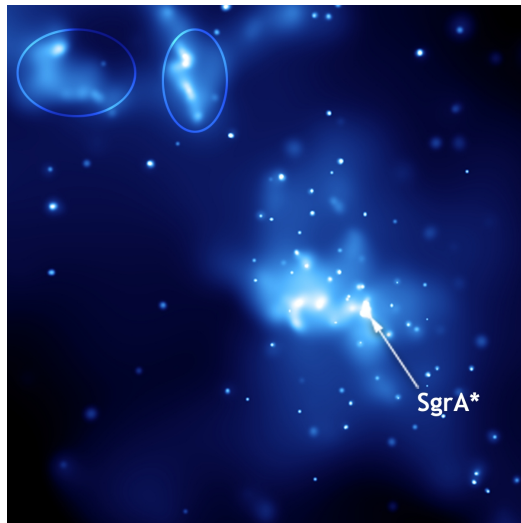
The black hole in our backyard: the challenge of observing Sgr A*

- Some facts about Sgr A*:
 - Located in the center of our galaxy.
 - Distance from Earth $\approx 7860\text{pc}$.
 - Estimated mass $4 \cdot 10^6 M_{\odot}$.
- ⇒ Nearby object of intermediate mass!



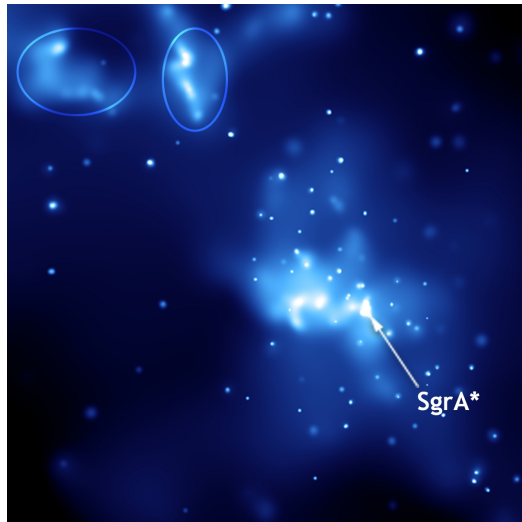
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 - ⇒ **Better candidate for direct observation?**



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 - ⇒ Better candidate for direct observation?
- Difficulties observing Sgr A*:
 - Matter moves at $\approx$ min timescales.
 - ⇒ Dynamic object, shorter integration time.
 - Interstellar medium scatters radio waves.
 - ⇒ Need to compensate for image distortion.



Summary

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⇒ Observations of dark objects are likely to have a bright future.