

Gravitational waves in teleparallel theories of gravity

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Center of Excellence "The Dark Side of the Universe"



Tartu cosmology workshop - 6. June 2018

- 1 Introduction
- 2 Waves in torsion gravity
- 3 Waves in non-metricity gravity
- 4 Conclusion

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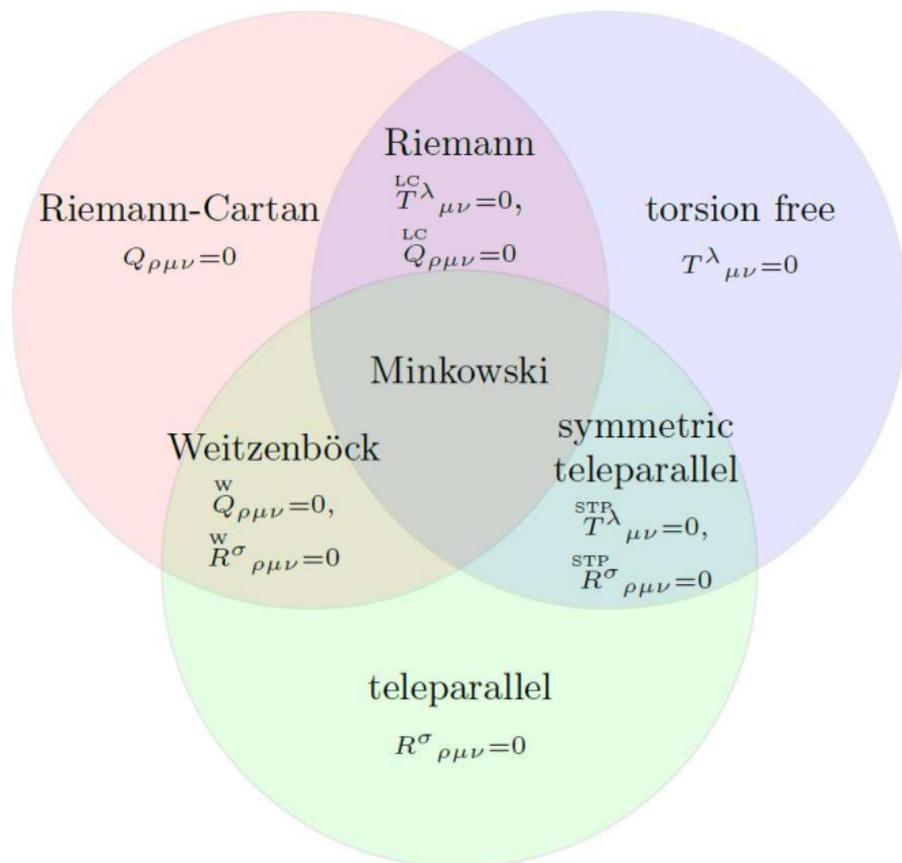
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- Gravity formulated as gauge theories.

Overview of geometries



- Complex double null basis of the tangent bundle:

$$l = \partial_t + \partial_z, \quad n = \frac{\partial_t - \partial_z}{2}, \quad m = \frac{\partial_x + i\partial_y}{\sqrt{2}}, \quad \bar{m} = \frac{\partial_x - i\partial_y}{\sqrt{2}}.$$

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- Consider plane null wave with $k_\mu = -\omega l_\mu$ and $u = t - z$:

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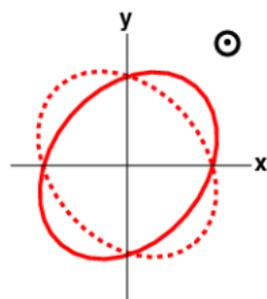
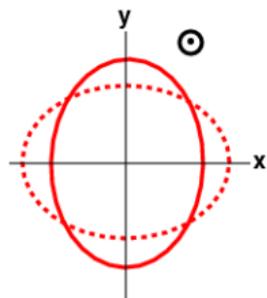
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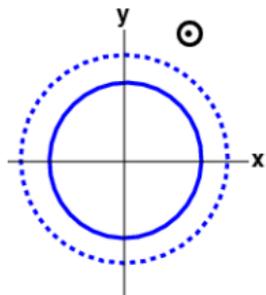
- Riemann tensor determined by “electric” components:

$$\begin{aligned} \Psi_2 &= -\frac{1}{6} R_{nlnl} = \frac{1}{12} \ddot{h}_{ll}, & \Psi_3 &= -\frac{1}{2} R_{nl\bar{m}\bar{m}} = \frac{1}{4} \ddot{h}_{l\bar{m}\bar{m}}, \\ \Psi_4 &= -R_{n\bar{m}n\bar{m}} = \frac{1}{2} \ddot{h}_{\bar{m}\bar{m}}, & \Phi_{22} &= -R_{nmn\bar{m}} = \frac{1}{2} \ddot{h}_{m\bar{m}}. \end{aligned}$$

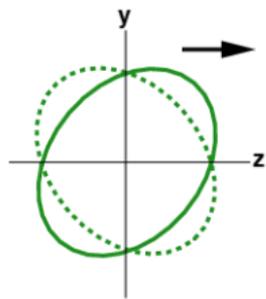
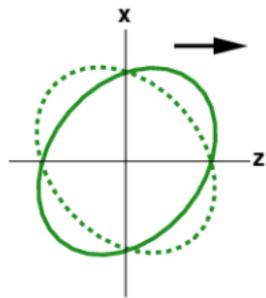
Polarisations of gravitational waves



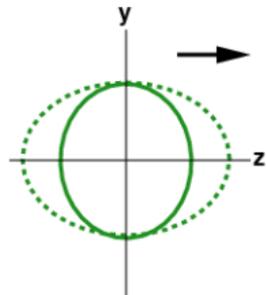
$\Psi_4, \bar{\Psi}_4$



Φ_{22}



$\Psi_3, \bar{\Psi}_3$



Ψ_2

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 - Flat spin connection $\dot{\omega}^a{}_b = \dot{\omega}^a{}_{b\mu} dx^\mu$.

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- Derived quantities:

- Frame field $e_a = e_a{}^\mu \partial_\mu$ with $\iota_{e_a} \theta^b = \delta_a^b$.
- Metric $g_{\mu\nu} = \eta_{ab} \theta^a{}_\mu \theta^b{}_\nu$.
- Volume form $\theta d^4x = \theta^0 \wedge \theta^1 \wedge \theta^2 \wedge \theta^3$.
- Levi-Civita connection

$$\overset{\circ}{\omega}_{ab} = -\frac{1}{2} (\iota_{e_b} \iota_{e_c} d\theta_a + \iota_{e_c} \iota_{e_a} d\theta_b - \iota_{e_a} \iota_{e_b} d\theta_c) \theta^c.$$

- Torsion $T^a = d\theta^a + \dot{\omega}^a{}_b \wedge \theta^b$.

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- Gauge fixing

- Perform local Lorentz transformation:

$$\theta'^a = \Lambda^a{}_b \theta^b, \quad \dot{\omega}'^a{}_b = \Lambda^a{}_c \dot{\omega}^c{}_d \Lambda_b{}^d + \Lambda^a{}_c d\Lambda_b{}^c.$$

⇒ Weitzenböck gauge: set $\dot{\omega}^a{}_b \equiv 0$.

Most general action and corresponding field equations

- Most general action:

$$S = \frac{1}{2\kappa^2} \int d^4x e (c_1 T^{\mu\nu\rho} T_{\mu\nu\rho} + c_2 T^{\mu\nu\rho} T_{\rho\nu\mu} + c_3 T^\mu{}_{\mu\rho} T^\nu{}^{\nu\rho}) .$$

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- Field tensors:

- Symmetric perturbation part:

$$F^{\mu\rho\sigma} = (2c_1 + c_2) (\partial^\sigma \phi^{\mu\rho} - \partial^\rho \phi^{\mu\sigma}) + c_3 [(\partial^\sigma \phi^\alpha{}_\alpha - \partial_\alpha \phi^{\alpha\sigma}) \eta^{\mu\rho} - (\partial^\rho \phi^\alpha{}_\alpha - \partial_\alpha \phi^{\alpha\rho}) \eta^{\mu\sigma}] .$$

- Antisymmetric perturbation part:

$$B^{\mu\rho\sigma} = (2c_1 - c_2) (\partial^\sigma a^{\mu\rho} - \partial^\rho a^{\mu\sigma}) + (2c_2 + c_3) \partial^\mu a^{\sigma\rho} .$$

Field equations expressed in Newman-Penrose basis

$$0 = E_{nn} = (2c_1 + c_2 + c_3)\partial_n^2\phi_{nl} + 2c_3\phi_{m\bar{m}} + (2c_1 + c_2 + c_3)\partial_n^2 a_{nl},$$

$$0 = E_{mn} = (2c_1 + c_2)\partial_n^2\phi_{ml} + (2c_1 - c_2)\partial_n^2 a_{ml},$$

$$0 = E_{\bar{m}n} = (2c_1 + c_2)\partial_n^2\phi_{\bar{m}l} + (2c_1 - c_2)\partial_n^2 a_{\bar{m}l},$$

$$0 = E_{nm} = -c_3\partial_n^2\phi_{lm} - (2c_2 + c_3)\partial_n^2 a_{lm},$$

$$0 = E_{n\bar{m}} = -c_3\partial_n^2\phi_{l\bar{m}} - (2c_2 + c_3)\partial_n^2 a_{l\bar{m}}$$

$$0 = E_{m\bar{m}} = -c_3\partial_n^2\phi_{ll},$$

$$0 = E_{ln} = (2c_1 + c_2)\partial_n^2\phi_{ll}, .$$

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 - Volume form $\sqrt{-\det g}d^4x$.
 - Levi-Civita connection

$$\overset{\circ}{\Gamma}{}^{\rho}{}_{\mu\nu} = \frac{1}{2}g^{\rho\sigma}(\partial_{\mu}g_{\sigma\nu} + \partial_{\nu}g_{\mu\sigma} - \partial_{\sigma}g_{\mu\nu}).$$

- Non-metricity $Q_{\rho\mu\nu} = \overset{\times}{\nabla}{}_{\rho}g_{\mu\nu}$.

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- Gauge fixing

- Perform local coordinate transformation:

$$g'_{\mu\nu} = \frac{\partial x^{\alpha}}{\partial x'^{\mu}} \frac{\partial x^{\beta}}{\partial x'^{\nu}} g_{\alpha\beta}, \quad \overset{\times}{\Gamma}'{}^{\rho}{}_{\mu\nu} = \frac{\partial x^{\alpha}}{\partial x'^{\mu}} \frac{\partial x^{\beta}}{\partial x'^{\nu}} \frac{\partial x'^{\rho}}{\partial x^{\gamma}} \overset{\times}{\Gamma}{}^{\gamma}{}_{\alpha\beta} + \frac{\partial^2 x^{\alpha}}{\partial x'^{\mu} \partial x'^{\nu}} \frac{\partial x'^{\rho}}{\partial x^{\alpha}}.$$

⇒ Coincident gauge: set $\overset{\times}{\Gamma}{}^{\rho}{}_{\mu\nu} \equiv 0 \Rightarrow Q_{\rho\mu\nu} = \partial_{\rho}g_{\mu\nu}$.

Most general action and corresponding field equations

- Most general action:

$$S = - \int d^4x \frac{\sqrt{-g}}{2} \left[c_1 Q^\alpha{}_{\mu\nu} + c_2 Q_{(\mu}{}^\alpha{}_{\nu)} + c_3 Q^\alpha g_{\mu\nu} + c_4 \delta_{(\mu}^\alpha \tilde{Q}_{\nu)} + \frac{c_5}{2} \left(\tilde{Q}^\alpha g_{\mu\nu} + \delta_{(\mu}^\alpha Q_{\nu)} \right) \right]$$

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- Linearized field equations:

$$0 = 2c_1 \eta^{\alpha\sigma} \partial_\alpha \partial_\sigma h_{\mu\nu} + c_2 \eta^{\alpha\sigma} (\partial_\alpha \partial_\mu h_{\sigma\nu} + \partial_\alpha \partial_\nu h_{\sigma\mu}) + 2c_3 \eta_{\mu\nu} \eta^{\tau\omega} \eta^{\alpha\sigma} \partial_\alpha \partial_\sigma h_{\tau\omega} \\ + c_4 \eta^{\omega\sigma} (\partial_\mu \partial_\omega h_{\nu\sigma} + \partial_\nu \partial_\omega h_{\mu\sigma}) + c_5 \eta_{\mu\nu} \eta^{\omega\gamma} \eta^{\alpha\sigma} \partial_\alpha \partial_\omega h_{\sigma\gamma} + c_5 \eta^{\omega\sigma} \partial_\mu \partial_\nu h_{\omega\sigma}.$$

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- Terms involving c_1 and c_3 do not contribute for a null wave $\square h_{\mu\nu} = 0$.

Field equations expressed in Newman-Penrose basis

$$0 = E_{nn} = -2(c_2 \ddot{h}_{ln} + c_4 \ddot{h}_{nl} + c_5 \ddot{h}_{nl} - c_5 \ddot{h}_{m\bar{m}}),$$

$$0 = E_{mn} = E_{nm} = -(c_2 + c_4) \ddot{h}_{lm},$$

$$0 = E_{\bar{m}n} = E_{n\bar{m}} = -(c_2 + c_4) \ddot{h}_{l\bar{m}},$$

$$0 = E_{m\bar{m}} = E_{\bar{m}m} = c_5 \ddot{h}_{ll},$$

$$0 = E_{nl} = E_{ln} = -(c_2 + c_4) \ddot{h}_{ll}.$$

Gravitational wave polarisations

$$c_2 = \sin \theta \cos \phi$$

$$c_4 = \sin \theta \sin \phi$$

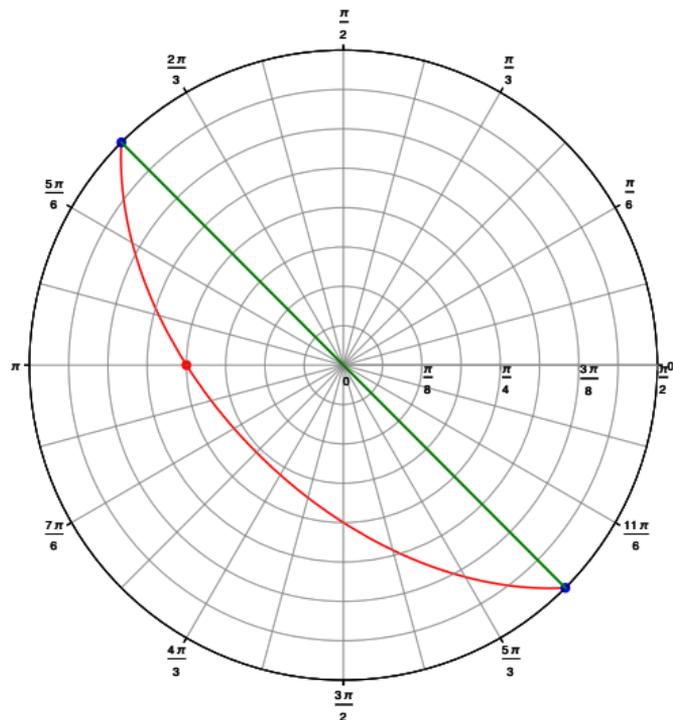
$$c_5 = \cos \theta$$

■ N_2

□ N_3

■ III_5

■ II_6



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- Results:
 - Gravitational waves propagate at the speed of light (not shown in this talk).
 - Polarisation classes N_2 , N_3 , III_5 , II_6 : tensor modes always exist, maybe more.

Teleparallel gravity workshop



June 25-29, 2018 - Tartu, Estonia

<http://hexagon.fi.tartu.ee/~telegrav2018/>

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- Estonian Research Council:



- European Regional Development Fund:



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